TABLE 2.-Vapor pressures at pyrheliometric stations on days when solar radiation intensities were measured.

Washington, D, C.			Madi	ison, W	is.	Lincoln, Nebr.			Santa	Fe, N. Mex.		
Dates.	s. m.	p. m.	Dates.	a. m.	p. m.	Dates.	a. m.	p. m .	Dates.	a. m.	p. m.	
1918. Aug. 12 14 15 16 21 23	mm. 19. 89 19. 23 12. 68 13. 13 11. 38 14. 60	19. 89 12. 68 15. 11 13. 61	13 19 22 24	mm, 14, 10 16, 20 11, 38 17, 96 15, 11 13, 13 13, 13	16. 79 10. 59 19. 23 12. 24 13. 13	13 26 29	mm. 13. 13 17. 37 11. 38 8. 48 7. 04	15. 11 7. 29 14. 10	1918. Aug. 1 2 3 5 9 10 13 15 17 19 23 28 27 29 30	mm. 7. 87 7. 87 7. 29 6. 27 10. 59 7. 87 8. 88 7. 87 9. 83 5. 79 9. 14 7. 29 7. 29 7. 87	7. 29 7. 87 10. 59 10. 59 7. 87 7. 29 6. 76	

TABLE 3.—Daily totals and departures of solar and sky radiation during August, 1918.

[Gram-calories per square centimeter of horizontal surface.]

	Da	ily tota	ls.	Dep	artures : normal.	from	Excess since f	or deficient	ficiency month.			
Day of month.	Wash- ing- ton.	Madi- son.	Lin- coin.	Wash- ing- ton,	Madi- son.	Lin- coln,	Wash- ing- ton.	Madi- son.	Lin- coln.			
Aug. 1	cal. 480 467 477 314 542 533 476 536 463 296	cal. 425 545 573 286 495 588 489 375 538 475	cal. 651 627 627 663 614 571 215 294 306 574	cal. - 3 - 15 - 4 - 166 63 56 2 65 - 5 - 169	cal 50 72 102 -183 28 124 27 - 84 81 21	cal. 115 94 96 135 88 48 -304 -222 -206 65	cal 3 - 18 - 22 - 188 - 125 - 69 - 67 - 2 - 7 - 176	cal 50 22 124 - 59 - 31 93 120 36 117 138	cal. 115 209 305 440 528 576 272 50 —156 — 91			
11	383 503 440 503 609 575 126 84 525 543	258 552 509 573 411 77 121 499 593 429	415 613 421 287 447 521 315 520 525 360	- 79 44 - 16 50 159 128 -318 -356 88 109	193 104 64 130 29 361 314 67 163	- 90 111 - 67 -208 - 44 33 -169 39 57 -115	- 255 - 211 - 227 - 177 - 18 110 - 208 - 564 - 476 - 367	- 55 49 113 243 214 -147 -461 -394 -231 -229	-181 - 70 -137 -345 -389 -356 -525 -486 -429 -544			
Decade depe	rture		1 :••••	;;	·	! ;	_ 191	-367	-453			
21	593 437 500 574 500 400 412 89 372 422 338	365 532 243 558 504 565 478 283 527 297 548	590 405 389 564 474 599 521 358 542 142 614	162 9 75 152 80 - 18 - 4 -325 - 40 12 - 70	- 60 110 -177 141 90 154 70 -122 125 -102 152	118 - 64 - 77 101 14 131 66 - 95 92 -306 168	- 205 - 196 - 121 31 111 93 89 - 236 - 276 - 264 - 334	-289 -179 -356 -215 -125 29 99 - 23 102 0 152	-426 -490 -567 -466 -452 -321 -255 -350 -258 -564 -396			
Decade depe	arture	I	l 	il 	J	! •••••	+ 33	+381	+148			
Excess or deficien	-1,266 - 1.3	+938 +1.0	+276 +0.2									

VOLCANIC ERUPTIONS AND SOLAR RADIATION INTENSITIES.

By HERBERT H. KIMBALL, Professor of Meteorology. [Dated: Washington, D. C., Sept. 13, 1918.]

In previous papers 1 curves have been presented showing variations in the annual mean values of solar radiation intensity at a few stations, for the period 1883-1912. These have been quoted by different writers,2

and it now seems advisable to bring them up to date, since the last curve published ends in the midst of the depression of 1912-13, following the eruption of Katmai Volcano, Alaska, in June, 1912.

In Table 1 there is given for each station the month and year of the beginning and ending, and reference to a footnote gives the character, of the data utilized in constructing this new curve. The data for some of the stations is not complete. For example, at Simla, India, radiation measurements are rarely obtained in July or August. At Mount Weather, Va., the records were frag-mentary until May, 1911, and the same is true of the early part of the records at Madison, Wis., and Santa Fe, N. Mex. The records of these two latter stations were not employed previous to 1914, except in determining the monthly normals for these stations.

TABLE 1.—Duration and character of solar radiation intensity records.

Year.	Montpellier, France.	Lausanne, Switzer- land.	Paviovsk, Russia.	Warsaw, Poland.	Washington, D. C.	Simla, India.	Paris, France.	Mount Weather, Va.	Madison, Wis.	Santa Fe, N. Mex.	Lincoln, Nebr.	Total number of stations.
1882 1883 1884 1885 1886 1886 1890 1890 1891 1892 1893 1894 1895 1896 1990 1900 1901 1902 1903 1904 1905 1906 1907 1909 1910 1910 1911 1912 1913 1914 1916 1918	Dec.	Jan. X X X X X X Aug.	Sept + + + + + + + + + + + + + + + + + + +	Jan. + + + + + + + + + + + + + + + + + + +	June OCCOO Aug.	Oct. XXXXXXXDec.	Jan. + + + + + Dec.	Sept.	<u>පූ</u> රගගටලිලි	(jet)	jaoo ja	111112112211233333333334666655545444

X—monthly means of noon solar radiation intensities.

+—monthly maxima of solar radiation intensities.

—monthly means of solar radiation intensities with the sun at zenith distance 60°.

(O)—data inclosed in curves employed only in completing the monthly normals of the respective stations.

For stations in the United States the monthly normal has been obtained by dividing the sum of all the radiation intensities measured in the months of a given name by the total number of measurements. For the other stations it has been necessary to take the average of the monthly means for the monthly normal.

The monthly means or the monthly maxima for each station have next been expressed as a percentage of either the monthly normal or the average monthly values, and the mean of the percentages for all the stations determined for each month. These monthly means have been

smoothed by the formula $\frac{a+2b+c}{a}$, where b is the mean

for the month in question, and a and c are the means for the preceding and following months, respectively.

¹ Kimball, Herbert H. Bulletin of the Mount Weather Observatory, 3:111; 5:301.
2 Humphreys, W. J. Volcanic dust and other factors in the production of climatic changes, and their possible relation to the ice ages. Bulletin of the Mount Weather Observatory, 6:1-34.

4 bbot, C. G., & Fowle, F. E. Volcanoes and climate. Annals of the Astrophysical Observatory of the Smithsonian Institution, 3:211-229.

The smoothed percentages thus obtained have been

plotted as dots on figure 1.

Reference to Table 1 shows that for the period December, 1882-August, 1892, data were available for a single station only, and from September, 1892, to December, 1895, from two stations only. Is consequence the monthly values are much more scattering previous to 1896, and especially previous to 1893, than subsequent to these years.

Furthermore, all the radiation measurements previous to 1905 were made at noon only. It follows from the Bouguer equation 3 that the percentage of depletion due to diminished transmissibility of the atmosphere, or the percentage increase due to increased atmospheric transmission, would be greater during the winter months, when the sun's zenithal distance is at a maximum, than in the summer months, when it is at a minimum. There is evidence of this annual variation in the radiation intensities, particularly in the early part of the curve.

obtained from stations in Europe. In consequence, the data are not so representative of conditions in the Northern Hemisphere as a whole, but are more influenced by local atmospheric conditions. Aside from the depressions noted, the smoothed monthly radiation intensities show remarkably uniform values. It is because the values obtained during the great depressions have been included in the means that the intensities measured under normal conditions generally fall between 102 and 104 per cent.

The following are the sources of the solar radiation

data that have been utilized:

Montpellier, France. Eon, L. J. Intensités calorifiques de la radiation solaire et durée de l'insolation a Montpellier. (Année 1900 et résumé de la série 1883-1900.) (In Bulletin météorologique du Départment de l'Hérault, année 1900. Montpellier. 1901.)

LAUSANNE, SWITZERLAND. Bührer, C. & Dufour, Henri. Résumé des observations actinométriques de l'année 1903. (Arch. sci. phys. nat., Genève. 4 pér. Tome 19. 1905. p. 388-390.) Dufour, Henri. L'insolation en Suisse. 1 partie. (Arch. sci. phys. nat., Genève. 4 pér. Tome 16. 1903. p. 417-434.) Dufour, Henri &

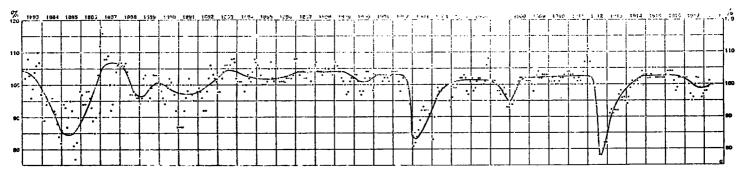


Fig. 1,-Monthly averages of solar radiation intensity at earth's surface, expressed as percentages of the monthly normals.

This has been taken into account in drawing a free-hand curve through the plotted values.

As previously shown,4 the three great depressions in the curve of figure 1 follow the great eruptions of Krakatoa in the summer of 1883; of Pelée, Santa Maria, and Colima, in 1902, and of Katmai in June, 1912. In each case there has been complete recovery—in the case of Krakatoa, after 34 months, of Pelée after 17 months, and of Katmai, after 21 months. There are minor depressions in 1888-89 and in 1907-8 that have not been accounted for definitely, and also one in 1916-17 that may be due to the fact that since the outbreak of the European war it has been necessary to substitute data from stations in the United States for that previously

Bührer, C. L'insolation en Suisse. 2 partie. (Arch. sci. phys. nat., Genève. 4 pér. Tome 16. p. 526-540.)

Pavlovsk, Russia. Savinov, S. I. [The maximum of solar radiation at Pavlovsk since 1892. The decrease in the intensity during the second half of 1912]. [Russian.] Bull., Ac. Imp. des. sci., St. Pétersbourg, 6 me sér., 1913, No. 12, p. 707-720.

WARSAW, POLAND. Gorczynski, Ladislav. Valeurs pyrhéliométriques et les sommes d'insolation à Varsovie pendant la période 1901-

1913. Varsovie, 1914.
Washington, D. C. Bulletin of the Mount Weather Observatory, 3:85-96; 5:182, 302, 303; Monthly Weather Review, 1914, 42:648; 1915, 43:112. [Monthly thereafter.]
Simia, India. Indian Monthly weather review, 1906-15.

PARIS (PARC St. MAUR), FRANCE. La Nature (Paris), Supplément, années 35-42, 1907-1914.

MOUNT WEATHER, VA. Bulletin of the Mount Weather Observatory, 5:303-311; 6:215-219. Monthly Weather Review, 1914, 42:139, 310, 520. Madison, Wis. Bulletin of the Mount Weather Observatory, 5:173-183. Monthly Weather Review, 1916, 44:3, 9-12. [Monthly thereafter.]

SANTA FE, N. MEX. Monthly Weather Review, 1915, 43:439-443.

591. [Monthly thereafter.] LINCOLN, NEBR. MON Monthly Weather Review, 1916, 44:4-8. [Monthly thereafter.]

 $^{^{3}}$ I=I₀ a^m, where I₀ is the radiation intensity outside the atmosphere, a is the transmission coefficient of the atmosphere, or the proportional part of the radiation transmitted when the sun is in the zenith, m is the length of the path of the sun's rays through the atmosphere in terms of the length of the path for zenithal sun, and I is the solar radiation intensity measured at the surface of the earth. This equation is strictly true for homogeneous rays only geneous rays only.

⁴ Kimball, H. H. loc. cit. Humphreys W. J. loc. cit.